

PHY 402 Atomic and Molecular Physics Instructor: Sebastian Wüster, IISER Bhopal, 2018

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2.4 Many-electron atoms

• Already cannot solve the two electron problem exactly. The variational approach also becomes too hard for large numbers of electrons. Now, the central field approximation will be very useful.

Many-electron (N) Schrödinger equation: $\hat{H}\psi(q_1, q_2, ..., q_N) = E\psi(q_1, q_2, ..., q_N)$ $q_k = \{m_{sk}, \mathbf{r}_k\} \text{ (spin and position variables)}$ $\hat{H} = \sum_{i=1}^N \left(-\frac{1}{2} \nabla_{\mathbf{r}_i}^2 - \frac{Z}{r_i} \right) + \sum_{i< j=1}^N \frac{1}{r_{ij}} \text{ (in atomic units)}$ (2.99)

- Eq. (2.4) constitutes a PDE, for a wavefunction with 3N spatial plus 2^N discrete co-ordinates \implies very impossible to solve directly.
- This problem generically arises in quantum-many-body-physics.
- Unlike two electron case, $\hat{H}' = \sum_{i < j=1}^{N} 1/r_{ij}$ generally is <u>not small</u> due to many terms in the sum.

2.4.1 Central field approximation

Now the independent-particle model and central field approximation become indispensable:

Nucleus Q=Z-e.

left: When considering a selected electron, the most important effect of other (N-1) electrons is collectively screening the nuclear charge, which however still leaves the potential for the Nth electron spherically symmetric (to a good approximation).

Similar to our treatment of Helium, we write for this radial potential felt by electron number i

$$V(\mathbf{r}_i) = -\frac{Z}{|\mathbf{r}_i|} + S(|\mathbf{r}_i|)$$
 (Now screening is *r*-dependent). (2.100)

We now re-write Eq.(2.99), ignoring spin for now

$$\hat{H}_{c} = \sum_{i=1}^{N} \left(-\frac{1}{2} \nabla_{\mathbf{r}_{i}}^{2} + V(\mathbf{r}_{i}) \right) = \sum_{i=1}^{N} \hat{h}_{i}$$
(2.101)

$$\hat{H}' = \sum_{i(2.102)$$

such that $\hat{H} = \hat{H}_c + \hat{H}'$ (just re-writing).

- \hat{H}_c is <u>central field Hamiltonian</u>.
- Expect eigenfunctions of \hat{H}_c to be better than those of $\hat{H}_o = \sum_{i=1}^N \left(-\frac{1}{2} \nabla_{\mathbf{r}_i}^2 Z/r_i \right).$
- Challenge is to find out suitable $S(\mathbf{r}_i)$ or $V(\mathbf{r}_i)$.

Central field Schrödinger equation

$$\ddot{H}_c \psi(\mathbf{r}_1, \mathbf{r}_2, ..., \mathbf{r}_N) = E_c \psi(\mathbf{r}_1, \mathbf{r}_2, ..., \mathbf{r}_N).$$
(2.103)

can be solved using

$$\psi(\mathbf{r}_1, \mathbf{r}_2, ..., \mathbf{r}_N) = u_{a_1}(\mathbf{r}_1) u_{a_2}(\mathbf{r}_2) ... u_{a_N}(\mathbf{r}_N); \ a_i \equiv \{n_i l_i m_i\}$$
(2.104)

where,

$$\left(-\frac{1}{2}\nabla_{\mathbf{r}_{i}}^{2}+V(\mathbf{r}_{i})\right)u_{a_{i}}(\mathbf{r}_{i})=E_{i}u_{a_{i}}(\mathbf{r}_{i}), \ E_{c}=\sum_{i}E_{i}.$$
(2.105)

The functions $u_{n_i m_i l_i}(\mathbf{r}_i)$ are called <u>central-field orbitals</u>.

• Excercise: Show directly that (2.104) is a solution for (2.103).

• To actually find the central field orbitals, as for hydrogen states, we write $u_{nlm}(\mathbf{r}) = R_{nl}(r)Y_{lm}(\theta,\phi)$ with Y_{lm} exact same as for H (spherical harmonics), but R_{nl} determined from the radial equation:

$$-\frac{1}{2}\left(\frac{d^2}{dr^2} + \frac{2}{r}\frac{d}{dr} - \frac{l(l+1)}{r^2}\right)R_{nl}(r) + \underbrace{V(r)}_{\neq}R_{nl}(r) = E_{nl}R_{nl}(r).$$
(2.106)
potential used for Hydrogen
due to screening

- Now we can use the following iterative procedure:
 - (i) Solve (2.106) with some trial $V(\mathbf{r}_i)$ to determine the $u_{a_i}(\mathbf{r}_i)$.
 - (ii) Use the obtained $u_{a_i}(\mathbf{r}_i)$ to infer the electron charge distributions.
 - (iii) Use the charge distribution to infer the screened potential $V(\mathbf{r}_i)$ with usual electro-statics methods.
 - (iv) Loop back to step (i) using the updated screened potential $V(\mathbf{r}_i)$.
 - (v) Repeat iteration until nothing changes any more.
- We will see one method that formalises this later.
- However we can learn a lot without doing the actual iteration, from two known limits:

$$V(r) \to -\frac{Z}{r} \text{ (for } r \mapsto 0\text{)},$$
 (2.107)

here the electron is closer to the nucleus than all others, so no screening can take place, and

$$V(r) \to -\frac{Z - (N - 1)}{r} \text{ (for } r \mapsto \infty), \qquad (2.108)$$

here all N - 1 other electrons are closer to the nucleus and screen its charge as much as possible.

2.4.2 Spin and Pauli-exclusion principle

• Product states $\psi(\mathbf{r}_1, ..., \mathbf{r}_N) = u_{a_1}(\mathbf{r}_1) ... u_{a_N}(\mathbf{r}_N)$ do not in general satisfy the anti-symmetry requirements of Eq. (1.62) for fermions, and we have not yet included spin $(\mathbf{r}_N \to q_N, \text{ see Eq.}(2.99))$.

Introduce spin-orbitals:

$$\underbrace{u_{nlm_lm_s}}_{\equiv \alpha}(q) = u_{nlm_l}(\mathbf{r}) | \chi \rangle_{m_s}$$
$$= R_{nl}(r) Y(\theta, \phi) | \chi \rangle_{m_s}$$
(2.109)

with

$$|\chi\rangle_{+\frac{1}{2}} = |\uparrow\rangle, |\chi\rangle_{-\frac{1}{2}} = |\downarrow\rangle.$$

In terms of these, a valid fermionic $N\text{-}\mathrm{electron}$ total wavefunction is given by a Slater-determinant

$$\psi_{c}(q_{1},...,q_{N}) = \frac{1}{\sqrt{N!}} \begin{vmatrix} u_{\alpha}(q_{1}) & u_{\beta}(q_{1}) & \cdots & u_{\nu}(q_{1}) \\ u_{\alpha}(q_{2}) & \ddots & & \vdots \\ \vdots & & \ddots & \vdots \\ u_{\alpha}(q_{N}) & \cdots & \cdots & u_{\nu}(q_{N}) \end{vmatrix}$$
(2.110)

Reminder: For matrix $A = \{a_{ij}\}$, the determinant $det A = |A| = \sum_{\substack{\sigma \in S_N \\ \text{Permutations}}} \operatorname{sgn}(\sigma) \prod_{i=1}^N a_{i\sigma_i}$.

- We see that this enforces antisymmetry, since determinant changes sign when we exchange two rows (*math course*).
- Also, if for any indices $\alpha = \beta \implies det = 0$. (Since, it also changes sign if we exchange two columns. This means there can never be two electrons in the same spin-orbital.)
- Set of electron quantum numbers $[\alpha, \beta, ..., \nu]$ in (2.110) is called electron-configuration.
- Let us define also

the total angular momentum of all electrons
$$\mathbf{J}_{tot} = \sum_{i=1}^{N} \mathbf{J}_i$$

the total orbital angular momentum of electrons $\mathbf{L}_{tot} = \sum_{i=1}^{N} \mathbf{L}_i$ (2.111)
the total spin of electrons $\mathbf{S}_{tot} = \sum_{i=1}^{N} \mathbf{S}_i$

• Note that $[H_c, \mathbf{L}_{tot}] = 0$, $[H_c, S_{tot}] = 0$ for the central field Hamiltonian H_c in Eq.(2.101).

 \implies We can in principle write all many electron eigenstates also as eigenstates of \hat{L}_{tot} , \hat{S}_{tot} . However, note, the Slater determinant (2.110) is not yet in general an eigenstate of \hat{L}_{tot} , \hat{S}_{tot} .

Atomic terms A state of multi-electron atoms with well defined quantum numbers for the total angular momenta J for \mathbf{J}_{tot} , L for \mathbf{L}_{tot} and S for \mathbf{S}_{tot} is called <u>term</u>, denoted with a term symbol

$$^{2S+1}L_J,$$
 (2.112)

c.f. Example p. 24.

2.4.3 Ground state energetic ordering and periodic table

• So far we have not really <u>solved</u> the many-electron-atom problem since we did not yet calculate $S(\mathbf{r})$. But, most essential properties of all atoms, incorporated into the periodic table, can be understood already now, based on some $V(\mathbf{r})$ with properties (2.107)-(2.108).

Angular momentum versus screening (N = Z atoms):

Revisit Eq.(2.106):

$$-\frac{1}{2}\left(\frac{d^2}{dr^2} + \frac{2}{r}\frac{d}{dr} - \frac{l(l+1)}{r^2}\right)R_{nl}(r) + \underline{V(r)}R_{nl}(r) = E_{nl}R_{nl}(r).$$
(2.113)

We can again combine the underlined pieces into an effective potential $V_{\text{eff}} = V(r) + \frac{l(l+1)}{2r^2}$ that includes the centrifugal barrier due to angular momentum (as in section 1.2.3).



- The centrifugal potential $l(l+1)/2r^2$ pushes the wavefunction further out (to larger r), the higher the value of l.
- All higher *l*-states thus feel the unscreened and stronger -Z/r potential less than lower *l* states thus energy typically increases with *l*.
- Energy E_i in Eq.(2.105) depends on n, l of all electrons (but not their m_l, m_s while we ignore fine and hyperfine structure).
 ⇒ electrons with the same n, l are said to belong to the same <u>sub-shell</u> and are also called equivalent electrons.

- For a given selection of the two numbers (n, l), there are 2(2l + 1) equivalent electrons (size of shell).
- To build the periodic table, we fill shells starting from low energies.

See table 8.3, Fig 8.1 of the book BJ, also follow https://ptable.com/

- -4s states fill before 3d (because 3d has so much centrifugal potential) \dagger .
- Filled sub-shells always have zero total spin and zero total angular momentum. This is because we have to fill all possible m_l , m_s and at the same time maintain fermionic total anti-symmetry.
- Still reminiscent of the situation in Hydrogen, the binding energy of the outer electron jumps up (becomes muss less negative) whenever $n \to n+1$, this gives <u>periodic recurrence of chemical</u> properties, mainly determined by outermost (valence) electron and how easily it is lost.
- The confusing middle part of the periodic table (transition metals) exists due to $3d \leftrightarrow 4s$ swap in the energy ordering (see † above) (This happens again for higher energy states $4d \leftrightarrow 5s, 4f \leftrightarrow 5d$).
- Note that all chemical variety in the world around us is due to the Pauli exclusion principle. Without it, any atom with varying nuclear charge Z would be just in a configuration $(1s)^Z$, thus all atoms would behave chemically very similar.

2.4.4 Approximation methods for many-electron systems

Thomas-Fermi theory: Assume electrons are (i) degenerate Fermi gas, (ii) numerous, (iii) can be treated in WKB/semi-classical approximation \implies Obtain electron charge density $\rho(\mathbf{r})$ and from that $V(\mathbf{r})$ (screened potential). Not very accurate, but interesting conceptually and useful in e.g. ultra-cold gases (also Bosons, see Chapter-V). Precursor to:

Density-functional theory: (Based on theorem by Hohenberg and Kohn)

For an N-electron system with ground-state wave function $\psi(\mathbf{r}_1, ..., \mathbf{r}_N)$, the ground-state electron density is

$$\rho(\mathbf{r}) = \int d^3 \mathbf{r}_2 \int d^3 \mathbf{r}_3 \cdots \int d^3 \mathbf{r}_N |\psi(\mathbf{r}, \mathbf{r}_2, ..., \mathbf{r}_N)|^2.$$
(2.114)

Then,

- (a) every observable is uniquely determined by $\rho(\mathbf{r})$.
- (b) there exists a functional $I[\rho]$ such that for a given potential $U(\mathbf{r})$ (felt by all e^{-}), ground state energy is minimum of

$$E_{\mathbf{o}}[\rho] = I[\rho] + \int d^3 \mathbf{r} U(\mathbf{r})\rho(\mathbf{r}), \text{ for } \int d^3 \mathbf{r} \rho(\mathbf{r}) = N.$$

- Useless unless we know $I[\rho]$. By now many "good" choices for $I[\rho]$ are available.
- Very successful method for quantum-chemistry/ material science.

Hartree-Fock method and the self-consistent field:

Idea: (i) Calculate $u_{nlm}(\mathbf{r})$ as in (2.105) using trial $V(\mathbf{r})$.

- (*ii*) Calculate electron charge density $\rho(\mathbf{r}) = -e \sum |u_{nlm}(\mathbf{r})|^2$, this gives screening $V'(\mathbf{r})$.
- (*iii*) Re-calculate $u'_{nlm}(\mathbf{r})$ with new $V'(\mathbf{r})$, iterate until converged.
- (*iv*) In practice, this is all embedded in a many-body variational method, so won't directly recognize ρ , V.

Now, <u>a sketch of derivation</u>:

Use natural splitting of Hamiltonian (2.99)

$$\hat{H}_1 = \sum_{i=1}^N \hat{h}_i, \quad \hat{h}_i = -\frac{\nabla_{\mathbf{r}_i}^2}{2} - \frac{Z}{r_i}, \quad \hat{H}_2 = \sum_{i< j=1}^N \frac{1}{r_{ij}}, \quad \hat{H} = \hat{H}_1 + \hat{H}_2$$

from variational principle (2.89), the true ground-state energy E_0 fulfills

$$E_0 \le E[\phi] = \langle \phi | \hat{H} | \phi \rangle, \qquad (2.115)$$

but now ϕ is a many-electron <u>Slater-determinant</u> as in (2.110). Let us re-write

$$\phi(q_1, ..., q_N) = \sqrt{N!} \mathscr{A} \phi_H(q_1, ..., q_N)$$
(2.116)

with <u>Hartree function</u>:

$$\phi_H(q_1, \dots, q_N) = u_\alpha(q_1) u_\beta(q_2) \dots u_\nu(q_N).$$
(2.117)

 $u_{\alpha}(q_i)$ etc. are again spin-orbitals, with form yet to be determined. We also used the anti-symmetrisation operator

$$\mathscr{A} \cdots = \frac{1}{N!} \sum_{\mathcal{P}} (-1)^{P} \mathcal{P}[\cdots]$$
(2.118)

Here the sum runs over all possible permutations \mathcal{P} of the set of integers $1 \cdots N$, P is the sign of the perturbation, and $\mathcal{P}[\cdots]$ tells us to permute coordinates that occur within the brackets according to \mathcal{P} , i.e. $u_{\alpha}(q_1)u_{\beta}(q_2) \rightarrow u_{\alpha}(q_{\mathcal{P}[1]})u_{\beta}(q_{\mathcal{P}[2]})$.

We now have to evaluate $E[\phi] = \langle \phi | \hat{H}_1 | \phi \rangle + \langle \phi | \hat{H}_2 | \phi \rangle$ for a Slater-determinant. We will use $[\hat{H}_i, \mathscr{A}] = 0$ (both \hat{H}_k are symmetric under all particle label swaps) and $\mathscr{A}^2 = \mathscr{A}$ (\mathscr{A} is a projector, if the expression in \cdots is already anti-symmetrised, anti-symmetrising it again does nothing).

Now:

$$\langle \phi | \hat{H}_{1} | \phi \rangle \stackrel{Eq. (2.116)}{=} N! \langle \phi_{H} | \mathscr{A} \hat{H}_{1} \mathscr{A} | \phi_{H} \rangle \stackrel{[\hat{H}_{i},\mathscr{A}]=0}{=} N! \langle \phi_{H} | \hat{H}_{1} \mathscr{A}^{2} | \phi_{H} \rangle$$
$$\stackrel{\mathscr{A} \text{ is }}{\underset{\text{projector}}{=}} N! \langle \phi_{H} | \hat{H}_{1} \mathscr{A} | \phi_{H} \rangle \stackrel{Eq. (2.118)}{=} \sum_{i=1}^{N} \sum_{\mathcal{P}} (-1)^{P} \langle \phi_{H} | \hat{h}_{i} \mathcal{P} | \phi_{H} \rangle$$

Now we can check that:

$$\langle \phi_H | \hat{h}_i \mathcal{P} | \phi_H \rangle = \begin{cases} \langle \phi_H | \hat{h}_i | \phi_H \rangle & ; \text{if } \mathcal{P} = \text{identity} \\ 0 & ; \text{else} \end{cases}$$
(2.119)

To see this explicitly expand the many-body scalar product as in the next step.

$$\begin{split} \langle \phi \, | \hat{H}_{1} | \, \phi \, \rangle &= \sum_{i=1}^{N} \langle \phi_{H} | \hat{h}_{i} | \phi_{H} \rangle \\ & \stackrel{Eq.(2.117)}{=} \sum_{i=1}^{N} \int dq_{1} \int dq_{2} \underbrace{\dots}_{\text{No}\,i} \int dq_{N} \left(u_{\alpha}^{*}(q_{1}) \, u_{\alpha}(q_{1}) \right) \left(u_{\kappa}^{*}(q_{i-1}) \, u_{\kappa}(q_{i-1}) \right) \left(u_{\eta}^{*}(q_{i+1}) \, u_{\eta}(q_{i+1}) \right) \\ & \left(u_{\nu}^{*}(q_{N}) \, u_{\nu}(q_{N}) \right) \left[\int dq_{i} \left(u_{\lambda}^{*}(q_{i}) \, \hat{h}_{i} u_{\lambda}(q_{i}) \right) \right] \\ & \stackrel{\text{all orbitals}}{\underset{\text{orthonormal}}{=}} \sum_{i=1}^{N} \langle u_{\lambda_{i}}(q_{i}) | \hat{h}_{i} | u_{\lambda_{i}}(q_{i}) \rangle \equiv \sum_{i=1}^{N} I_{\lambda_{i}}; \quad \text{using an index vector } \vec{\lambda} = [\alpha, \beta, ..., \nu]. \end{split}$$

The I_{λ_i} can be thought of as non-interacting energy of electron number *i* being in the spin-orbital $u_{\lambda_i}(q_i)$.

Similarly:

$$\langle \phi | \hat{H}_{2} | \phi \rangle = \sum_{i < j} \sum_{\mathcal{P}} (-1)^{P} \underbrace{\left\langle \phi_{H} \middle| \frac{1}{r_{ij}} \mathcal{P} \middle| \phi_{H} \right\rangle}_{\text{whenever P "touches" any index } \neq i, j \text{ this is } = 0$$

$$\underset{i < j}{\overset{\text{only allowed}}{\underset{permutation}{\underbrace{\sum_{i < j}}}} \sum_{i < j} \left\langle \phi_{H} \middle| \frac{1}{r_{ij}} (1 - \mathcal{P}_{ij}) \middle| \phi_{H} \right\rangle$$

$$\underset{i < j}{\overset{\text{as}}{\underset{above}{\underbrace{\sum_{\lambda, \mu}}}} \underbrace{\sum_{i < j}}_{\underset{(\text{Pairs})}{\underbrace{\left\langle u_{\lambda}(q_{i}) u_{\mu}(q_{j}) \middle| \frac{1}{r_{ij}} \middle| u_{\lambda}(q_{i}) u_{\mu}(q_{j}) \right\rangle}}_{\equiv \mathcal{J}_{\lambda\mu}} - \underbrace{\left\langle u_{\lambda}(q_{i}) u_{\mu}(q_{j}) \middle| \frac{1}{r_{ij}} \middle| u_{\mu}(q_{i}) u_{\lambda}(q_{j}) \right\rangle}_{\underset{i < \lambda\mu}{\underbrace{\operatorname{direct term, compare Eq. (2.97)}}}$$

In the second line, the only allowed permutations that don't give zero in the scalar product, is the identity (giving the 1), and \mathcal{P}_{ij} which flips $i \leftrightarrow j$.

• All together, we have now obtained a much simpler looking energy functional

$$E[\phi] = \sum_{\lambda} I_{\lambda} + \frac{1}{2} \sum_{\lambda \mu} \left[\mathcal{J}_{\lambda \mu} - \mathcal{K}_{\lambda \mu} \right].$$
(2.120)

That consists of single particle energies I_{λ} and interaction energies J and K. The 1/2 just avoids double counting of pairs.

• Now, in a more complex variational principle than used for Helium, we vary <u>all</u> the $u_{\lambda}(q_i)$ themselves:

$$\delta E - \sum_{\lambda} \underbrace{E_{\lambda}}_{\lambda} \delta \langle u_{\lambda} | u_{\lambda} \rangle = 0.$$
Lagrange-multiplier
(see math and book)
$$(2.121)$$

The Lagrange multipliers implement the constraint $\langle u_{\lambda} | u_{\lambda} \rangle = \delta_{\lambda\mu}$.

From variation $\delta u_{\lambda}(q_i)$ we obtain (see books)

Hartree-Fock equations:

$$\left[-\frac{1}{2} \nabla_{\mathbf{r}_{i}}^{2} - \frac{Z}{r_{i}} \right] u_{\lambda}(q_{i}) + \left[\sum_{\mu} \int dq_{j} \, u_{\mu}^{*}(q_{j}) \frac{1}{r_{ij}} u_{\mu}(q_{j}) \right] u_{\lambda}(q_{i}) - \left[\sum_{\mu} \int dq_{j} \, u_{\mu}^{*}(q_{j}) \frac{1}{r_{ij}} u_{\lambda}(q_{j}) \right] u_{\mu}(q_{i}) = E_{\lambda} u_{\lambda}(q_{i}).$$
(2.122)

- Set of integer-differential, coupled equations for N spin orbitals $u_{\lambda}(q_i)$.
- Lagrange-multipliers E_{λ} take role of energy eigenvalues.
- We can define a direct potential

$$V^{(d)}(q_i) = \sum_{\mu} \int dq_i u_{\mu}^*(q_i) \frac{1}{r_{ij}} u_{\mu}(q_i)$$

and exchange potential (operator)

$$V^{(ex)}(q_i)[f] = \sum_{\mu} \left[\int dq_i u^*_{\mu}(q_i) \frac{1}{r_{ij}} f(q_i) \right] u_{\mu}(q_i)$$

such that

$$\left[-\frac{1}{2}\nabla_{\mathbf{r}_{i}}^{2} - \underbrace{\frac{Z}{r_{i}} + V^{(d)} - V^{(ex)}}_{\equiv V(q_{i})}\right]u_{\lambda}(q_{i}) = E_{\lambda}u_{\lambda}(q_{i}).$$
(2.123)

Comparison with Eq.(2.103), $V(q_i)$ takes the role of <u>central field</u> here.

- For atoms with filled subshells, can show that $V(q_i)$ is indeed spherically symmetric. Otherwise not, but derivations are small.
- Solve (2.122) by <u>iteration</u>: Start with some trial solution $u_{\lambda}^{(0)}(q_i)$ (e.g. the one for Hydrogen or guessed $V(q_i)$). Then find solution of

$$\begin{bmatrix} -\frac{1}{2}\nabla_{\mathbf{r}_{i}}^{2} - \frac{Z}{r_{i}} \end{bmatrix} u_{\lambda}^{(1)}(q_{i}) + \begin{bmatrix} \sum_{\mu} \int dq_{j} \, u_{\mu}^{*(0)}(q_{j}) \frac{1}{r_{ij}} u_{\mu}^{(0)}(q_{j}) \end{bmatrix} u_{\lambda}^{(1)}(q_{i}) \\ - \begin{bmatrix} \sum_{\mu} \int dq_{j} \, u_{\mu}^{*(0)}(q_{j}) \frac{1}{r_{ij}} u_{\lambda}^{(0)}(q_{j}) \end{bmatrix} u_{\mu}^{(1)}(q_{i}) = E_{\lambda} u_{\lambda}^{(1)}(q_{i}). \quad (2.124)$$

etc. until $u_{\lambda}^{(n)}(q_i)$ is converged.



$2.4.5 \quad {\rm Corrections \ to \ the \ central \ field. \ L-S/J-J \ coupling}$

There are two important correction terms to the central field picture discussed so far:

$$\hat{H}_1 = \sum_{i < j=1}^N \frac{1}{r_{ij}} - \sum_{i=1}^N \left(\frac{Z}{r_i} + V(\mathbf{r}_i)\right), \qquad (2.126)$$

$$\hat{H}_2 = \sum_{i=1}^{N} \zeta_i(r) \,\hat{L}_i \cdot \hat{S}_i.$$
(2.127)

First is <u>non-central field</u> part of Hamiltonian, which we had listed as \hat{H}' in Eq. (2.102). Second spin-orbit coupling for each electron (c.f. Eq. (2.5)), the most important of relativistic corrections neglected in Eq. (2.99).

Which one is more important depends on Z.

L-S coupling (or Russell-Sanders coupling): : Warning: L-S coupling means sth. very different from the words "spin-orbit coupling"!! L-S coupling is used for small or intermediate Z, in this case

 $|\hat{H}_1| \gg |\hat{H}_2|$ (nontrivial to link this to "small Z").

- First consider only \hat{H}_1 . \hat{H} commutes with \mathbf{J}_{tot} , \mathbf{L}_{tot} , \mathbf{S}_{tot} , see section 2.4.2.
- Determining which quantum numbers J, L, S are possible for a given state is nontrivial due to Pauli-exclusion principle⁰. Energy level with given fixed values of L, S is called a <u>term</u>.

Hund's rules: For ground states:

- the term with the largest possible S for a given configuration has the lowest energy; energy then goes up as S goes down.
- for a given S, the term with maximum L has the lowest energy.
- So far we considered \hat{H}_1 only. Now we can add \hat{H}_2 as perturbation \rightarrow <u>Fine-structure</u>.
- Now each term with fixed L, S splits into separate J-components, which form a multiplett.
- <u>Landé interval rule</u>:

$$E(J) - E(J-1) = \overline{A} \times J$$
, within one multiplett, (2.128)

where \bar{A} is a constant. One finds $\bar{A} > 0$ for less than half filled shells, so that the groundstate has the <u>lowest</u> allowed J. However, $\bar{A} < 0$ for more than half filled shells, so that the ground-state then has the hightest allowed J.

⁰e.g. see section 4, helium ground-state can be only L = 0, not L = 1



J-J coupling: This takes place for large Z ($Z \approx 80$), then we have $|\hat{H}_2| \gg |\hat{H}_1|$.

- <u>First</u>: $\hat{H}_c + \hat{H}_2 = \sum_{i=1}^N \hat{h}_i \rightarrow$ for each electron separately couple spin and angular momentum to j_k , giving orbitals u_{nljm_j} (based on $\hat{J}_i = \hat{L}_i + \hat{S}_i$).
- <u>Second</u>: We only then consider $\hat{H}_c + \hat{H}_2 + \hat{H}_1$ and have to label our many-electron state with $\hat{J} = \sum_{i=1}^{N} \hat{J}_i$, only labelled by J.

