

Week 4

PHY 402 Atomic and Molecular Physics

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2.3 Two Electron Atoms

These include, e.g., $\underbrace{H^-}_{Z=1}$, $\underbrace{He}_{Z=2}$, $\underbrace{Li^+}_{Z=3}$

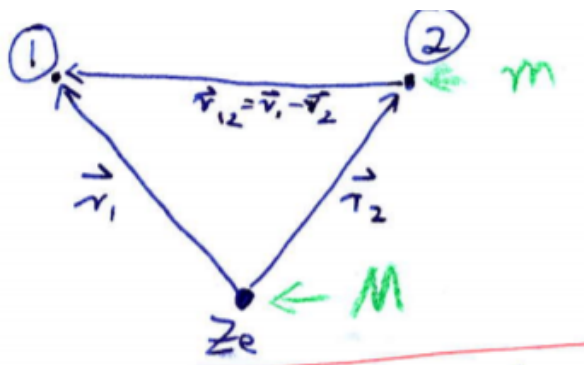
Important because they are the simplest atoms where we see the Pauli Exclusion Principle at work (see Eq. (1.62)) and can introduce essential approximation techniques. We cannot solve any two e^- problems exactly analytically.

2.3.1 The Schrödinger Equation for two-electron atoms

We need co-ordinates for both electrons now, see diagram below.

$$r_{12} = |\mathbf{r}_1 - \mathbf{r}_2| \quad (2.64)$$

$$r_1 = |\mathbf{r}_1| \quad (2.65)$$



In terms of these we write the

TISE (see Eq. (1.7)) for the Helium problem

$$\underbrace{\left[-\frac{\hbar^2}{2\mu} \nabla_{\mathbf{r}_1}^2 - \frac{\hbar^2}{2\mu} \nabla_{\mathbf{r}_2}^2 - \frac{\hbar^2}{M} \nabla_{\mathbf{r}_1} \cdot \nabla_{\mathbf{r}_2} - \frac{Ze^2}{(4\pi\epsilon_0)r_1} - \frac{Ze^2}{(4\pi\epsilon_0)r_2} + \frac{e^2}{(4\pi\epsilon_0)r_{12}} \right]}_{\equiv \hat{H}_{hel}} \psi(\mathbf{r}_1, \mathbf{r}_2) = E\psi(\mathbf{r}_1, \mathbf{r}_2) \quad (2.66)$$

- $\mu = \frac{m_e M}{m_e + M}$ is reduced mass of the electron, m_e electron mass, M nuclear mass. (We used $\mu = m_e$ for $M \rightarrow \infty$ in section 1.2.3)
- Mass polarisation term $\nabla_{\mathbf{r}_1} \cdot \nabla_{\mathbf{r}_2}$ comes from separation of centre of mass coordinate. It vanishes for $M \rightarrow \infty$.

From now on, use atomic units:

$$\hbar = 1 \quad \frac{1}{4\pi\epsilon_0} = 1 \quad e = 1 \quad \underbrace{m}_{\text{electron mass}} = 1 \quad (2.67)$$

(then hydrogen energy $E_{nlm} = -\frac{1}{2n^2}$ (see Eq. (1.43)) and $a_0 = 1$)

- The Hamiltonian is symmetric with respect to particle **position** interchange operator

$$P_{12} : \mathbf{r}_1 | \mathbf{r}_2 \longrightarrow \mathbf{r}_2 | \mathbf{r}_1 \quad (2.68)$$

\Rightarrow Also eigen-functions must have this symmetry [=be eigenfunction of P_{12}]

$$\Rightarrow \psi(\mathbf{r}_1, \mathbf{r}_2) = \pm \psi(\mathbf{r}_2, \mathbf{r}_1) \quad (2.69)$$

- Note, this is not the same as the complete particle exchange operation used in writing Bose/Fermi symmetries Eq. (1.62), where we have to swap all properties, not just position.

There are two types of two electron states

Para States, wave functions spatially symmetric:

$$\psi_+(\mathbf{r}_1, \mathbf{r}_2) = \psi_+(\mathbf{r}_2, \mathbf{r}_1) \quad (2.70)$$

Ortho States, wave functions spatially anti-symmetric:

$$\psi_-(\mathbf{r}_1, \mathbf{r}_2) = -\psi_-(\mathbf{r}_2, \mathbf{r}_1) \quad (2.71)$$

2.3.2 Spin Wave functions and Pauli Exclusion Principle

- Now we add electron spin into the picture, it becomes essential now, not just a small perturbation as for Hydrogen.
- e^- are fermions, total state must be anti-symmetric under $1 \leftrightarrow 2$
- Possible spin states for two electrons are: $|\uparrow\uparrow\rangle, |\uparrow\downarrow\rangle, |\downarrow\uparrow\rangle, |\downarrow\downarrow\rangle$, see section 1.2.2.
- Useful to move to coupled spin basis, where $\hat{S} = \hat{s}_1 + \hat{s}_2$, see also section 1.2.2.

We find

Spin pair states

$$|S = 0, m_S = 0\rangle = \frac{1}{\sqrt{2}} (|\uparrow\downarrow\rangle - |\downarrow\uparrow\rangle) \quad \begin{array}{l} \text{Anti-symmetric} \\ \text{spin singlet} \end{array} \quad (2.72)$$

$$|S = 1, m_S = -1\rangle = |\downarrow\downarrow\rangle \quad \text{Symmetric} \quad (2.73)$$

$$|S = 1, m_S = 0\rangle = \frac{1}{\sqrt{2}} (|\uparrow\downarrow\rangle + |\downarrow\uparrow\rangle) \quad \text{spin triplet} \quad (2.74)$$

$$|S = 1, m_S = 1\rangle = |\uparrow\uparrow\rangle \quad (2.75)$$

Solutions to the helium problem have to satisfy Eq. (2.70)-(2.71) (spatial symmetry) and Eq. (1.62) (total fermionic symmetry). Thus all allowed solutions can be written as

$$\psi(q_1, q_2) = \psi_+(\mathbf{r}_1, \mathbf{r}_2) \frac{1}{\sqrt{2}} (|\uparrow\downarrow\rangle - |\downarrow\uparrow\rangle) \quad (\text{Para}) \quad (2.76)$$

$$\psi(q_1, q_2) = \psi_-(\mathbf{r}_1, \mathbf{r}_2) \begin{cases} |\downarrow\downarrow\rangle \\ \frac{1}{\sqrt{2}} (|\uparrow\downarrow\rangle + |\downarrow\uparrow\rangle) \\ |\uparrow\uparrow\rangle \end{cases} \quad (\text{Ortho}) \quad (2.77)$$

- Despite not actually appearing explicitly in the Hamiltonian of (2.66), the total spin dictates which spatial symmetry the solution must have!

2.3.3 Approximate methods for 2 electron groundstate wavefunction

So far we only classified solutions according to spin and spatial symmetry, we yet have to actually determine the symmetric and anti-symmetric eigen functions $\psi_{\pm}(\mathbf{r}_1, \mathbf{r}_2)$

Lets split $\hat{H}_{hel} = \hat{H}_0 + \hat{H}'$ again, where (in atomic units, abbreviated a.u.)

$$\hat{H}_0 = -\frac{\nabla_{r_1}^2}{2} - \frac{\nabla_{r_2}^2}{2} - \frac{Z}{r_1} - \frac{Z}{r_2} \quad \hat{H}' = \frac{1}{r_{12}} \quad (2.78)$$

Note: Atomic units save a lot of writing!

We see that the unperturbed eigen-problem $\hat{H}_0\psi^{(0)}(\mathbf{r}_1, \mathbf{r}_2) = E^{(0)}\psi^{(0)}(\mathbf{r}_1, \mathbf{r}_2)$ is solved by the product Ansatz:

$$\psi^{(0)}(\mathbf{r}_1, \mathbf{r}_2) = \phi_{nlm}(\mathbf{r}_1)\phi_{n'l'm'}(\mathbf{r}_2), \quad (2.79)$$

$$E^{(0)} = E_{nlm} + E_{n'l'm'}, \quad (2.80)$$

with ϕ_{nlm} and E_{nlm} given by Hydrogen solutions ((1.35) and (1.43)).

- Note, that for every energy $E^{(0)}$, the wavefunctions with swapped indices: $\phi_{n'l'm'}(\mathbf{r}_1)\phi_{nlm}(\mathbf{r}_2)$ are an equally valid solution. This is called \Rightarrow exchange degeneracy.
- This allows us now to construct solutions satisfying Eq. (2.70)-(2.71), namely

Zero'th order approximation for Helium wavefunctions

This is also called independent particle model

$$\psi_{\pm}^{(0)}(\mathbf{r}_1, \mathbf{r}_2) = \frac{1}{\sqrt{2}} [\phi_{nlm}(\mathbf{r}_1)\phi_{n'l'm'}(\mathbf{r}_2) \pm \phi_{n'l'm'}(\mathbf{r}_1)\phi_{nlm}(\mathbf{r}_2)] \quad (2.81)$$

Where + = Para, and - = Ortho. Only for Para states $\{n', l', m'\} = \{n, l, m\}$ is possible, then $\psi_{+}^{(0)}(\mathbf{r}_1, \mathbf{r}_2) = \phi_{nlm}(\mathbf{r}_1)\phi_{nlm}(\mathbf{r}_2)$.

- Gives ground-state energy of Helium : $E_{100,100}^{(0)} = -\frac{Z^2}{2} \left(\frac{1}{1^2} + \frac{1}{1^2} \right) = -Z^2 \stackrel{\text{He}}{=} -4$
(Para only)

So far we have completely ignored electron-electron interactions in \hat{H}' , let's rectify this now: PerturbationTheory

Take into account \hat{H}' as in Eq. (1.50), focussing on the ground-state only for now:

$$E^{(1)} = \langle \psi_{\pm}^{(0)} | \hat{H}' | \psi_{\pm}^{(0)} \rangle \quad (2.82)$$

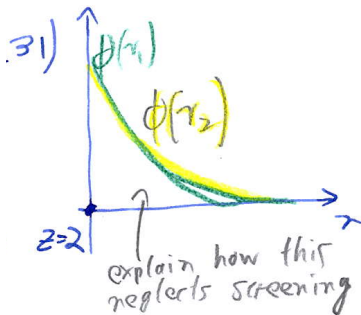
$$\stackrel{\text{Ground}}{=} \int d^3r_1 \int d^3r_2 \underbrace{|\phi_{100}(r_1)|^2}_{\text{Charge density at } \mathbf{r}_1} \frac{1}{|\mathbf{r}_1 - \mathbf{r}_2|} \underbrace{|\phi_{100}(r_2)|^2}_{\text{Charge density at } \mathbf{r}_2} \quad (2.83)$$

$$\stackrel{\text{See}}{\text{Book}} \dots + \frac{5}{8}Z \quad \text{in a.u.} \quad (2.84)$$

In the second line above we recognize the electro-static interaction energy of the charge densities due to electron 1 and 2.

- Now we have $E^{(0)} + E^{(1)} = -Z^2 + \frac{5}{8}Z = \underbrace{-2.75}_{\text{getting better}}$ ("exact" -2.904)

Variational Method



left: Sketch of helium electron wave-functions with (yellow) and without screening (green).

Look at

$$\psi_+^{(0)}(\mathbf{r}_1, \mathbf{r}_2) = \phi_{100}(\mathbf{r}_1) \phi_{100}(\mathbf{r}_2), \quad (2.85)$$

$$\phi_{100}(\mathbf{r}_1) \stackrel{\text{a.u.}}{\text{Eq. (1.35)}} \sqrt{\frac{Z^3}{\pi}} \exp[-Zr_1] \quad (2.86)$$

Now try to improve on this by introducing a screened (or effective) charge $Z \rightarrow Z_{\text{eff}}$. Let us take the trial state

$$\psi^{(0)}_{\text{trial}}(\mathbf{r}_1, \mathbf{r}_2) = \phi_{\text{trial}}(\mathbf{r}_1) \phi_{\text{trial}}(\mathbf{r}_2)$$

$$\phi_{\text{trial}}(\mathbf{r}) = \sqrt{\frac{Z_{\text{eff}}^3}{\pi}} \exp[-Z_{\text{eff}}r], \quad (2.87)$$

as guess for the two-electron wave function. We have to find the "best" value of Z_{eff} from variational principle

Variational Principle

- Energy functional

$$E[\phi] = \frac{\langle \phi | \hat{H} | \phi \rangle}{\langle \phi | \phi \rangle} \quad (2.88)$$

is extremal at an eigenstate ψ_n of \hat{H} , that means

$$\delta E[\psi_n] = 0. \quad (2.89)$$

- The groundstate energy $E_0 \leq E[\phi]$ for any trial state ϕ . Thus the minimal energy we can reach by changing the parameters in our trial state, will be closest to the true ground-state energy.

Variational derivatives:

- δE is a variation of the energy functional: Multiply Eq. (2.88) by $\langle \phi | \phi \rangle$ and take the total derivative wrt. the state

$$\delta E \langle \phi | \phi \rangle + E \langle \delta \phi | \phi \rangle + E \langle \phi | \delta \phi \rangle = \langle \phi | \hat{H} | \delta \phi \rangle + \langle \delta \phi | \hat{H} | \phi \rangle \quad (2.90)$$

(*"How does the energy $E[\phi]$ change for a small variation $\phi + \delta\phi$ around ϕ ?"*)

- Related to functional derivative $\frac{\delta E}{\delta \phi(x)}$ (see books)
- A special variation is $\delta E = \frac{\partial E}{\partial Z_{\text{eff}}} \delta Z_{\text{eff}} \stackrel{!}{=} 0$

To find $\frac{\partial E}{\partial Z_{\text{eff}}}$, we insert Eq. (2.87) into the energy functional Eq. (2.88):

$$E \left[\psi^{(0)}_{\text{trial}} \right] = \langle \psi^{(0)}_{\text{trial}} | \underbrace{\hat{T}_1 + \hat{T}_2 - \frac{Z}{r_1} - \frac{Z}{r_2}}_{\text{see QM1/ standard methods}} + \frac{1}{r_{12}} | \psi^{(0)}_{\text{trial}} \rangle = Z_{\text{eff}}^2 - 2Z Z_{\text{eff}} + \frac{5}{8} Z_{\text{eff}}. \quad (2.91)$$

- The term involving $1/r_{12}$ gives rise to the same integral as in the perturbation theory segment one page above.
- In the expression above, factors Z stem from the Hamiltonian, but factors of Z_{eff} from the trial function.

- $$\frac{\partial E}{\partial Z_{\text{eff}}} = 0 \Leftrightarrow Z_{\text{eff}} = Z - \frac{5}{16} \quad \text{effective charge reduced, as expected} \quad (2.92)$$

- Now ground state energy of Helium $E[\phi_{\text{trial}}] = -\left(Z - \frac{5}{16}\right)^2 \text{ a.u.} = -2.848 \text{ a.u.}$ (even closer to "exact" -2.904)

Screening in the Central Field approximation

Going back to independent particle model/ 0'th order P.T, we can rewrite Hamiltonians as indicated by the vertical arrow below:

| \hat{H}_0 | \hat{H}' |
|--|--|
| $-\frac{\nabla_{r_1}^2}{2} - \frac{\nabla_{r_2}^2}{2} - \frac{Z}{r_1} - \frac{Z}{r_2}$ | $\frac{1}{r_{12}}$ |
| | \Downarrow |
| $-\frac{\nabla_{r_1}^2}{2} - \frac{\nabla_{r_2}^2}{2} + V(r_1) + V(r_2)$ | $\frac{1}{r_{12}} - V(r_1) - V(r_2) - \frac{Z}{r_1} - \frac{Z}{r_2}$ |
| | (2.93) |

This just constitutes a re-allocation of what we call \hat{H}_0 and what we call \hat{H}' .

See that if we choose the **central field** below for our potential energy,

$$V(r) = -\frac{Z-S}{r} = -\frac{Z_{\text{eff}}}{r} \quad (2.94)$$

with $Z_{\text{eff}} = Z - \frac{5}{16}$ we obtain variational solution as eigenstates in 0'th order PT.²

The re-writing has effectively made \hat{H}' "smaller". S is the screening factor. ($S = \frac{5}{16} = 0.31$ for He)

- This central field concept will be even more useful for $N > 2$ electrons.

2.3.4 Excited States of Two electron atoms

Much of section 2.3.3 can be generalised to excited states.

Let us consider perturbation theory of:

$$\psi_{\pm}^{(0)}(\mathbf{r}_1, \mathbf{r}_2) = \frac{1}{\sqrt{2}} [\phi_{100}(\mathbf{r}_1) \phi_{nlm}(\mathbf{r}_2) \pm \phi_{nlm}(\mathbf{r}_1) \phi_{100}(\mathbf{r}_2)] \quad (2.95)$$

The unperturbed energy is $E^{(0)} = E_{nlm} + E_{100}$. For the perturbed results we find:

First Order excited state with interaction energy

$$E_{\pm}^{(1)} = J \pm K \quad + = \text{para}, - = \text{ortho}, (\text{ortho not possible for ground state}) \quad (2.96)$$

$$J = \int d^3\mathbf{r}_1 d^3\mathbf{r}_2 |\phi_{100}(\mathbf{r}_1)|^2 \frac{1}{r_{12}} |\phi_{nlm}(\mathbf{r}_2)|^2 \quad \text{Coulomb (direct) integral} \quad (2.97)$$

$$K = \int d^3\mathbf{r}_1 d^3\mathbf{r}_2 \phi_{100}^*(\mathbf{r}_1) \phi_{nlm}^*(\mathbf{r}_2) \frac{1}{r_{12}} \phi_{100}(\mathbf{r}_2) \phi_{nlm}(\mathbf{r}_1) \quad \text{Exchange integral} \quad (2.98)$$

- Using $\hat{\mathbf{S}}_1 \cdot \hat{\mathbf{S}}_2 = \frac{1}{2} \hat{\mathbf{S}}^2 - \frac{3}{4}$, we can express the energies as

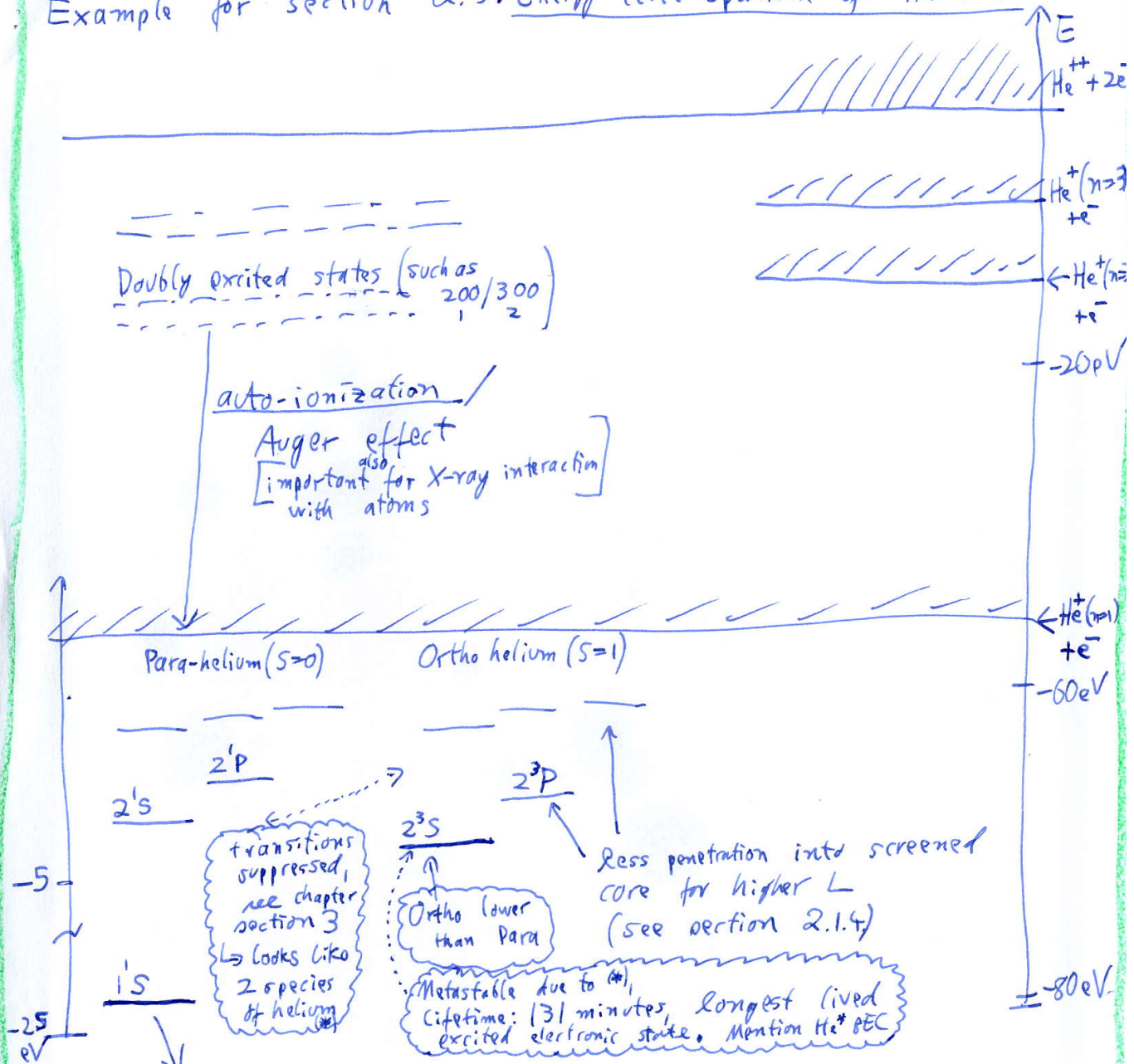
$E_{\pm}^{(1)} = J - \frac{1}{2} (1 + 4\hat{\mathbf{S}}_1 \cdot \hat{\mathbf{S}}_2) K \Rightarrow$ spin dependence of energy. Note that the energy became spin dependent because of the required allocation (2.76)-(2.77) of spatial symmetry to total electron spin.

- When evaluating integrals we find that $J > 0$ (see BJ book). It also turns out that J, K depend on n, l , hence ($J \rightarrow J_{nl}, K \rightarrow K_{nl}$).
- Usually also $K > 0 \Rightarrow$ Thus the ortho state ($S = 1$) has a lower energy.

All that we have learnt so far (and some more) enters the energy level diagram of helium on the next page:

²Finding eigenstates of the newly arranged \hat{H}_0 in (2.93) proceeds as we did for finding Eq. (2.81), except in the solution we have to replace everywhere $Z \rightarrow Z_{\text{eff}}$. Thus we obtain the solution (2.87).

Example for section 2.3: Energy level spectrum of Helium



Helium groundstate, ionization potential
 including fine structure etc...
 theory (fcl): $24.586812487292 \text{ eV} \pm 6 \cdot 10^{-6} \text{ eV}$
 expt. : $24.5868274589038 \text{ eV} \pm 1.8 \cdot 10^{-5} \text{ eV}$
 (QM seems to work)